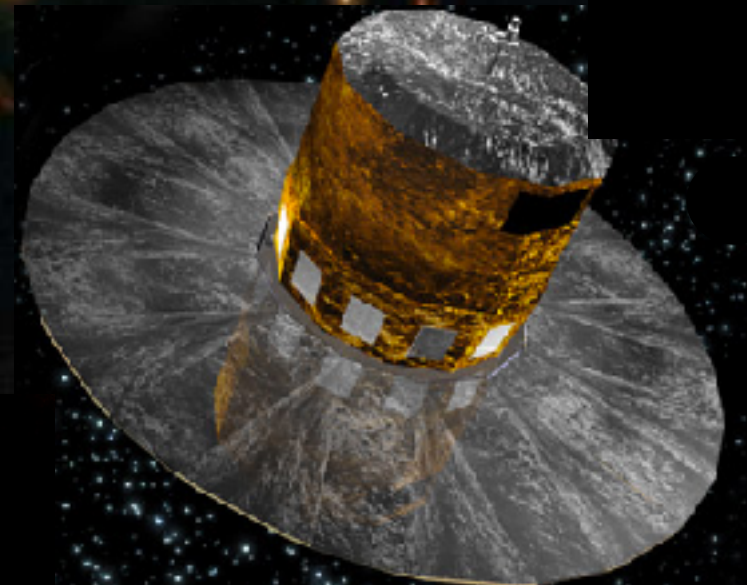


GRAVITATIONAL LENSES IN GAIA

L. Delchambre, F. Finet (Liège), L. Galluccio (Nice), A. Krone-Martins (Lisboa), JF Le Campion (Bordeaux), F. Mignard, E. Slesak (Nice), J. Surdej (Liège), R. Teixeira (São Paulo), O. Wertz (Liège)


Christine Ducourant - Alberto Krone-Martins



ORIGINS OF THE PROJECT



Firenze - CU4 meeting- 2013

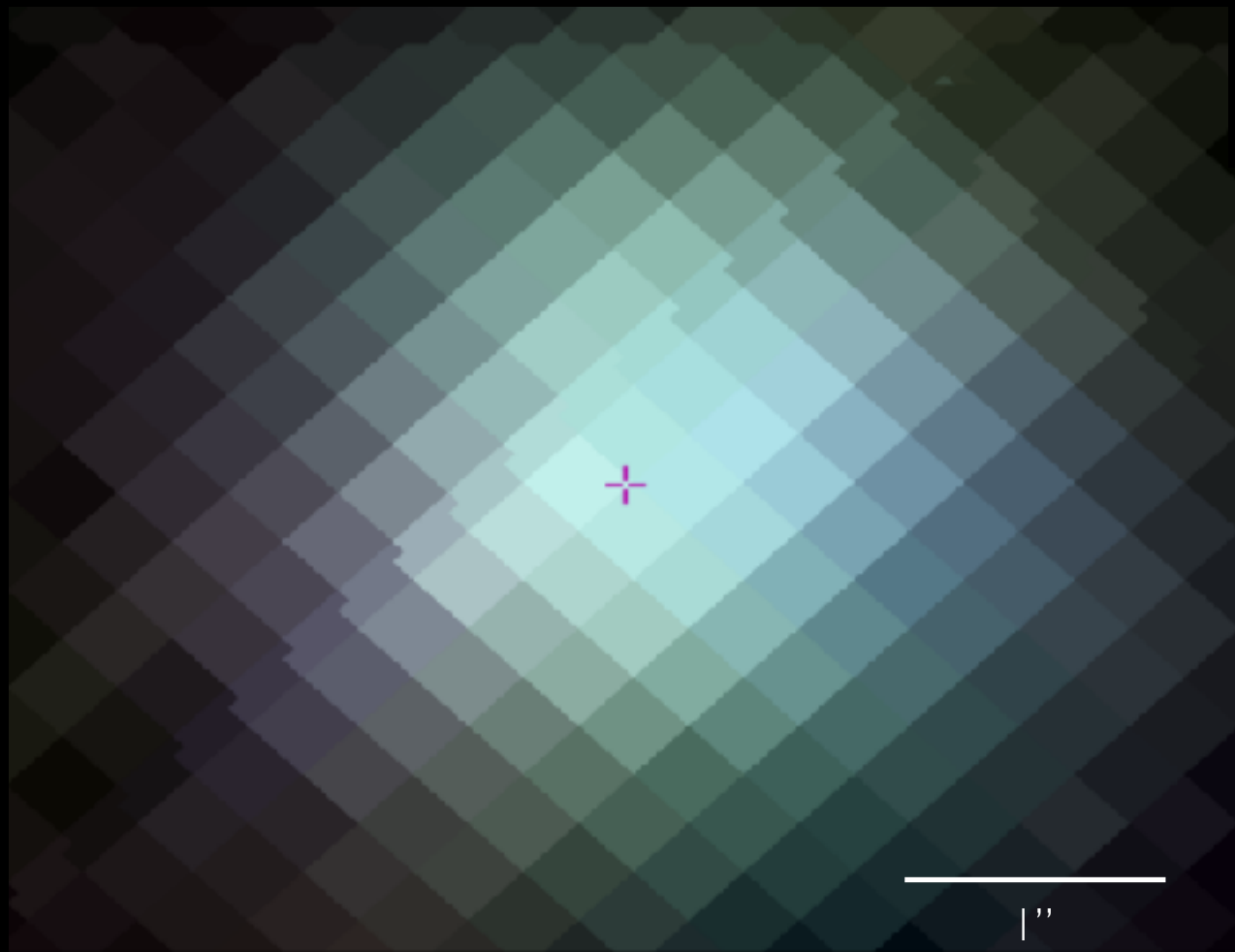


Gaia observes many extragalactic objects
(1 M Galaxies + 0.5 M QSOs)

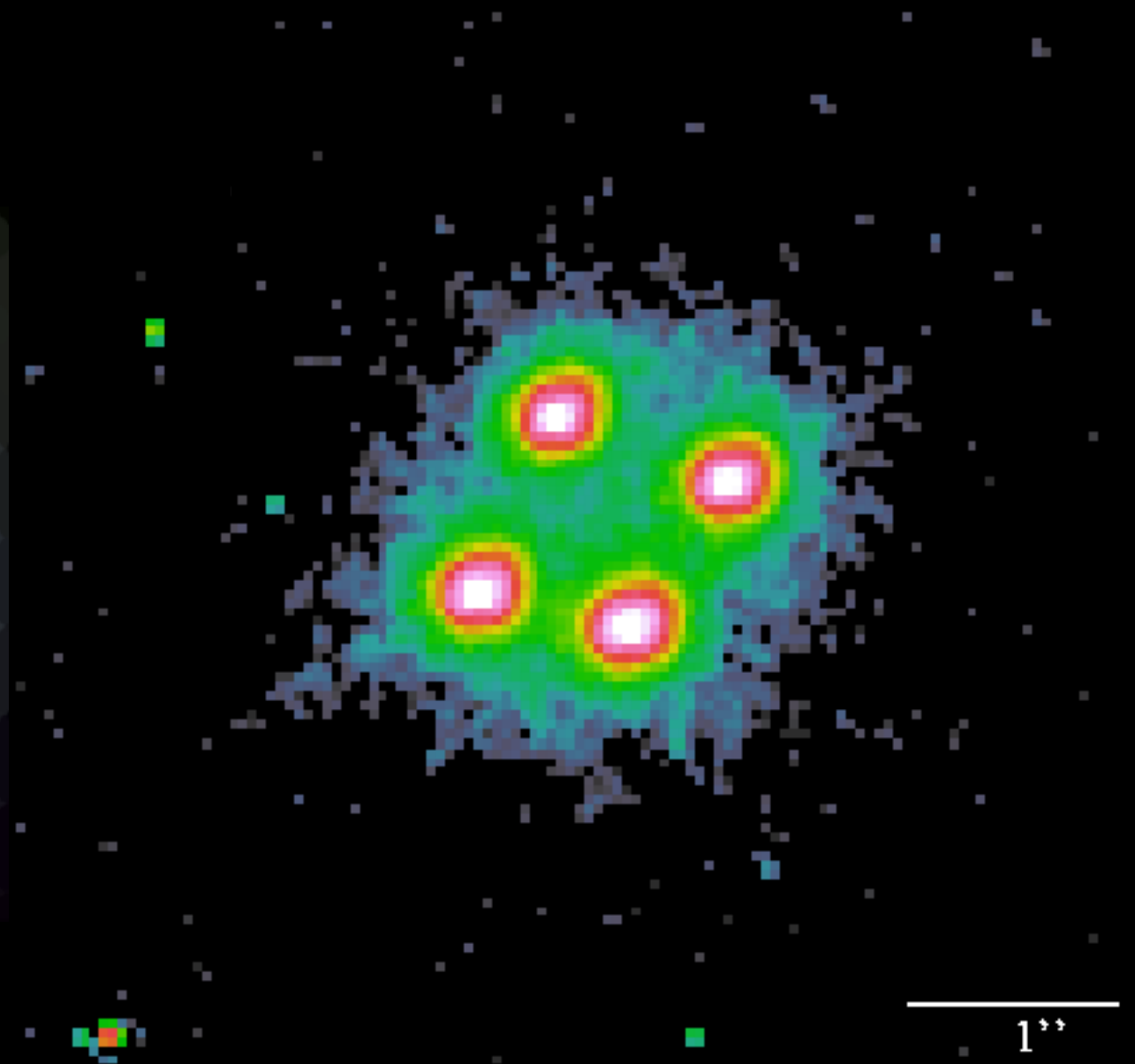
Although designed for the Milky Way
content, it is the
**first all-sky survey from space of
gravitational lenses (GL)**
with the acuity of **unprecedented
resolving power!**

Gravitational Lenses in Gaia

H1413+117
From Ground (SDSS9)



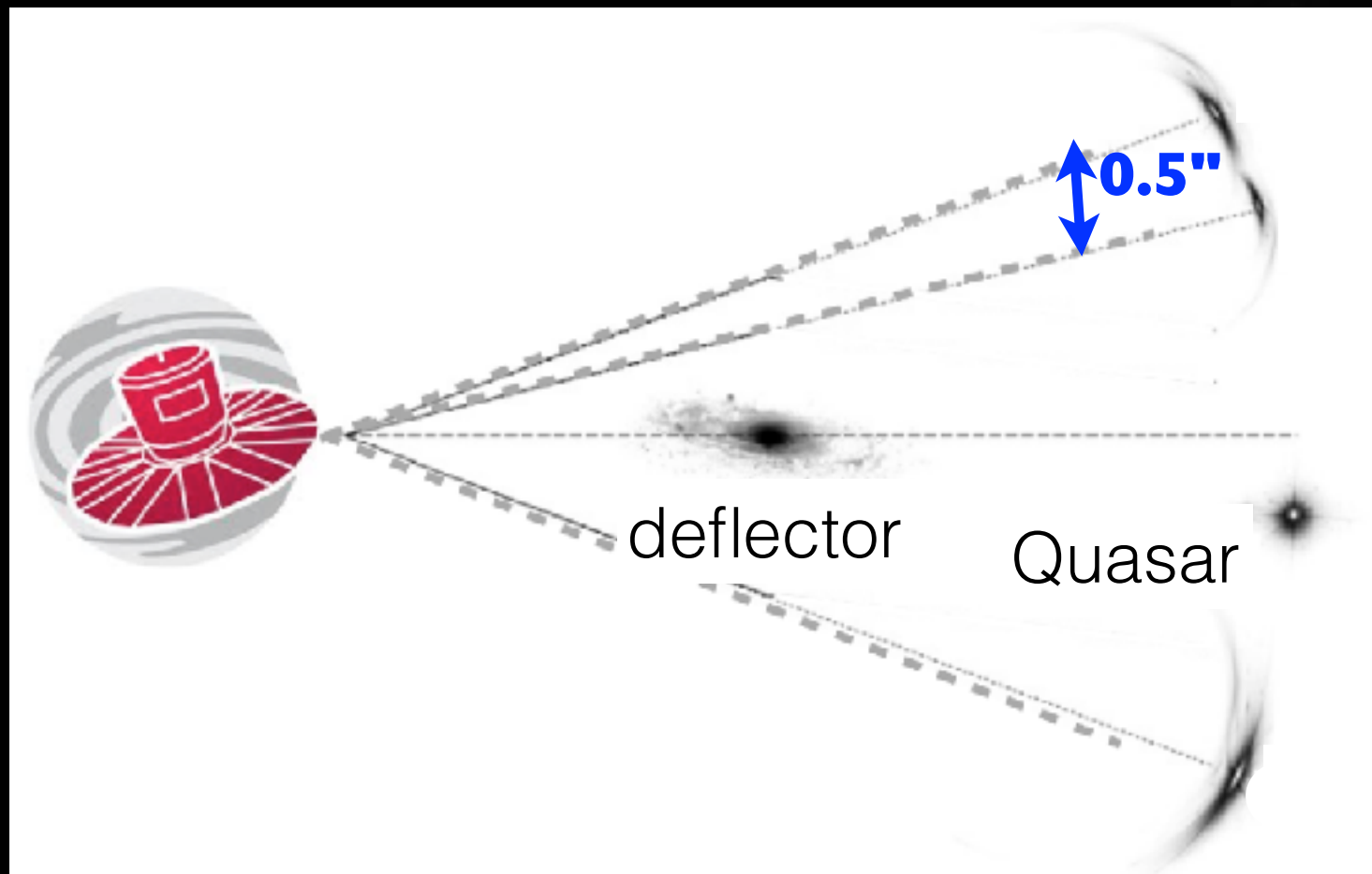
H1413+117
From space (HST)



Credit : Castles database, HST images

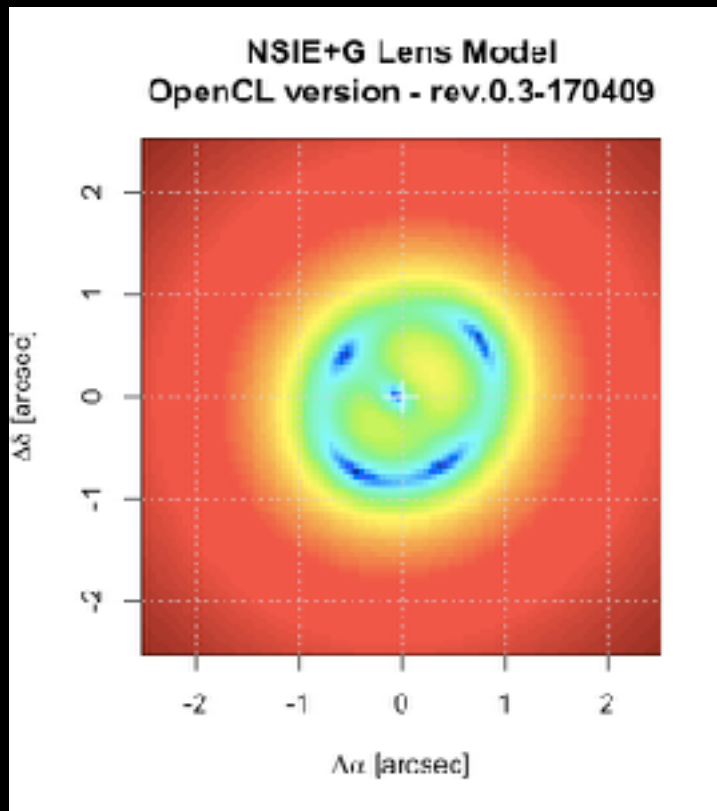
Gravitational Lenses in Gaia

- Depending on deflecting galaxy population **100 - ~1000 lensed quasars in Gaia**
- Most cases : 2 images, few 3 or 4 images, no arc!
- ~100 known today
- ➔ **Counts will bring independent constraints on cosmological parameters H_0 , Ω_0 , λ_0 , mass distribution of the galaxy lenses...**

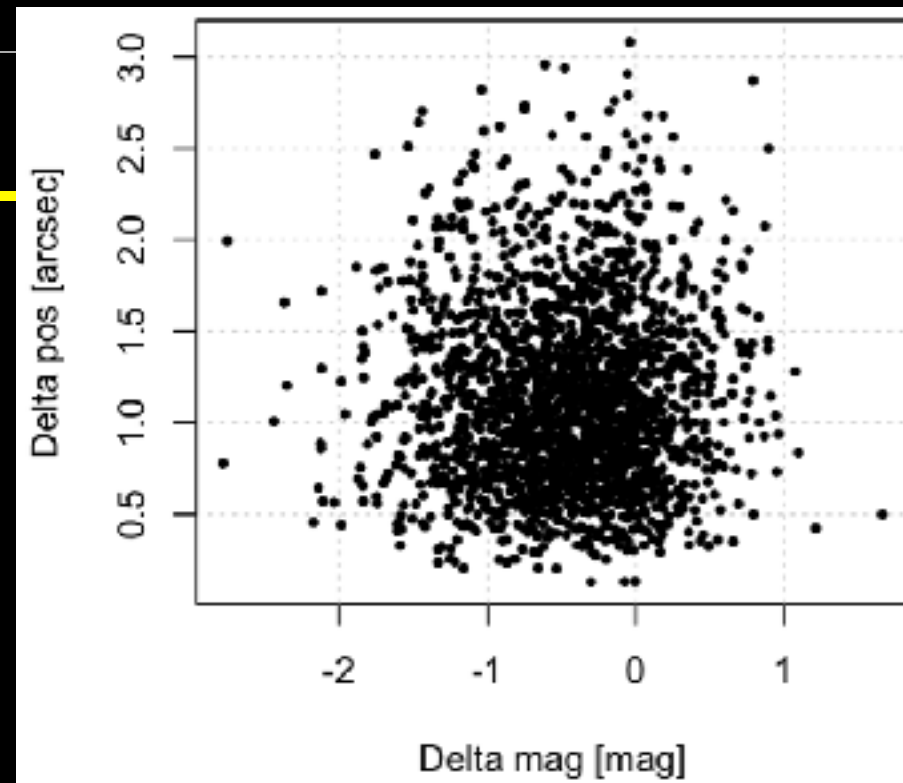


Detection of lenses, how ?

Specific ($\Delta\text{mag}, \Delta\text{pos}$)

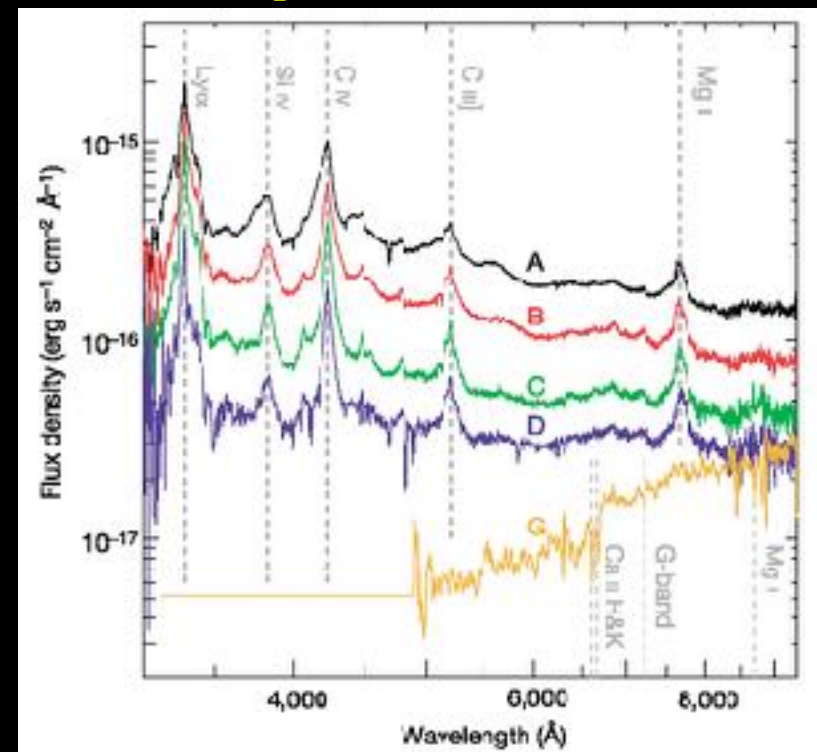


Krone-Martins 2017



Specific patterns

Same colour/spectra



Inada 2003



1 - Look for known lenses

Updated Database

DRI + Gaia observations (IDT/IDU)

2 - Search around known quasars

Automatic search for optical projections around quasars

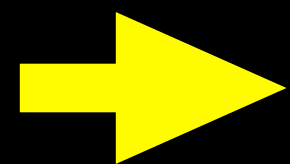
3 - Blind search in Gaia DRI

Search for lens configurations in DRI (Δ_{mag} , Δ_{pos})

Modeling of the detected candidate lenses

1 - Known Lenses : a new database

- ➔ **Castles** (<https://www.cfa.harvard.edu/castles/>),
 - HST images (same resolution as Gaia) , 100 lenses+ 18 bin quasars (<2006)
- ➔ **SQLS** (<http://www-utap.phys.s.u-tokyo.ac.jp/~sdss/sqls/lens.html>)
 - SDSS, 49 additional lenses (<2012) —> unresolved and resolved multi images
- ➔ **SDSS - III BOSS quasar lens survey**
 - 13 lensed quasars (2images) + 11 quasar pairs
- ➔ **others** (BELLS, DES, CASSOWARY,...)



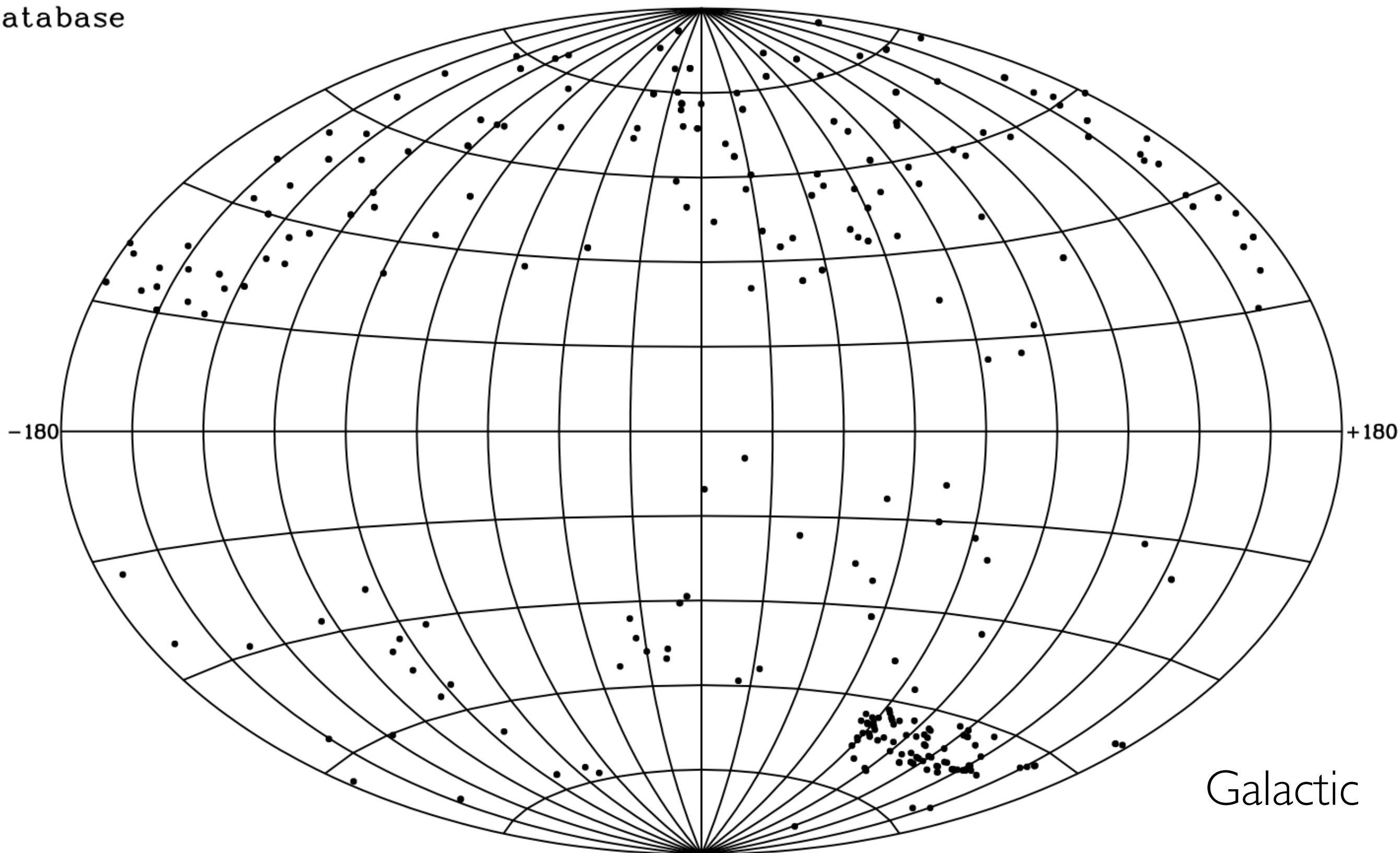
276 lenses/candidates

165 - with 1 component
90 - with 2 components
20 - with ≥ 3 components

Heterogeneous database with sometimes poor astrometry

Sky coverage and detection

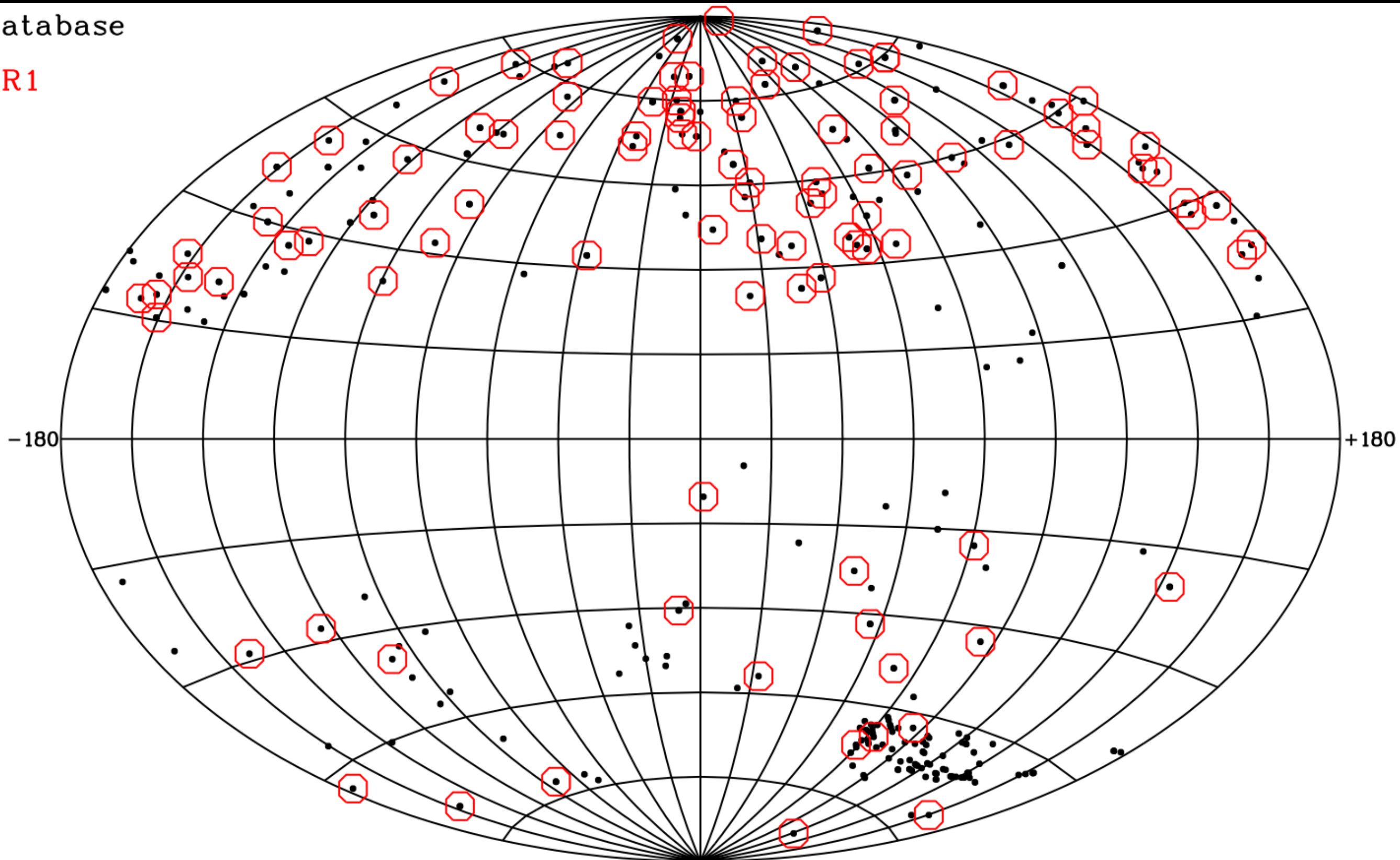
Database



Sky coverage and detection

Database

DR1

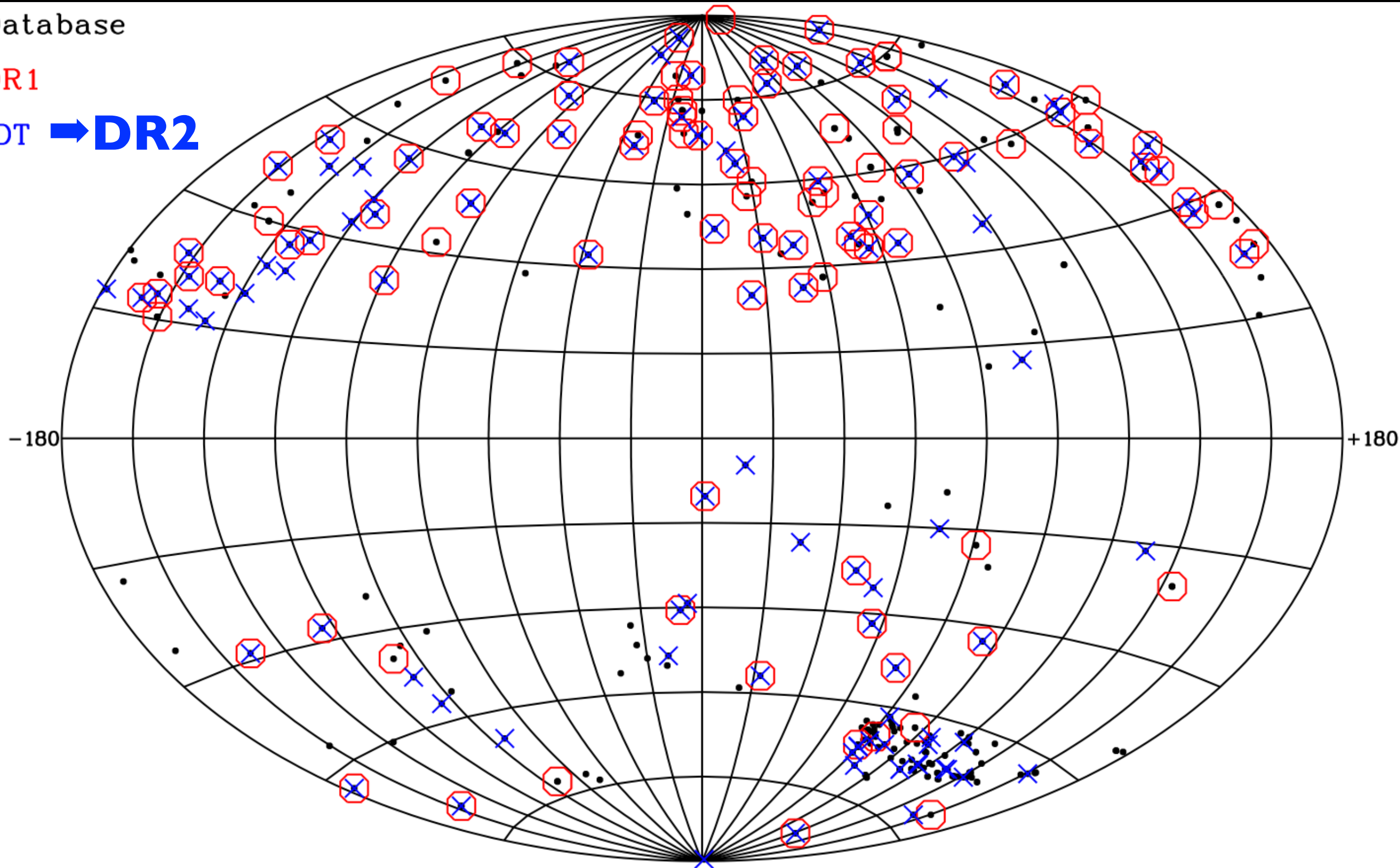


Sky coverage and detection

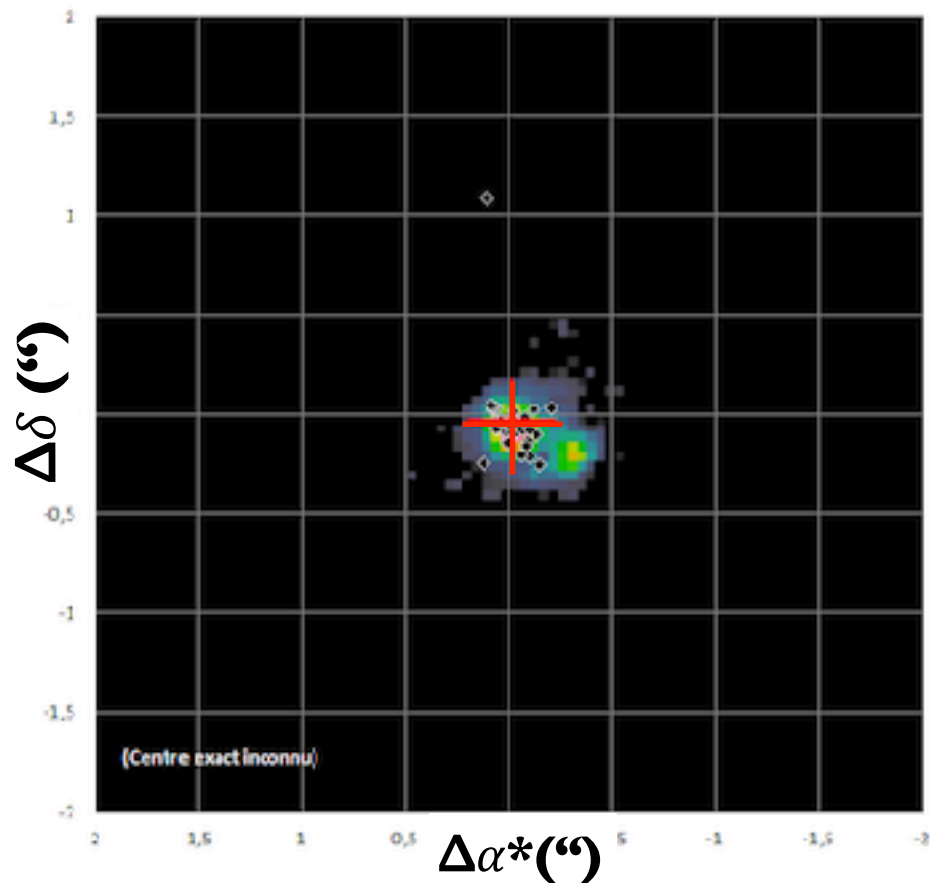
Database

DR1

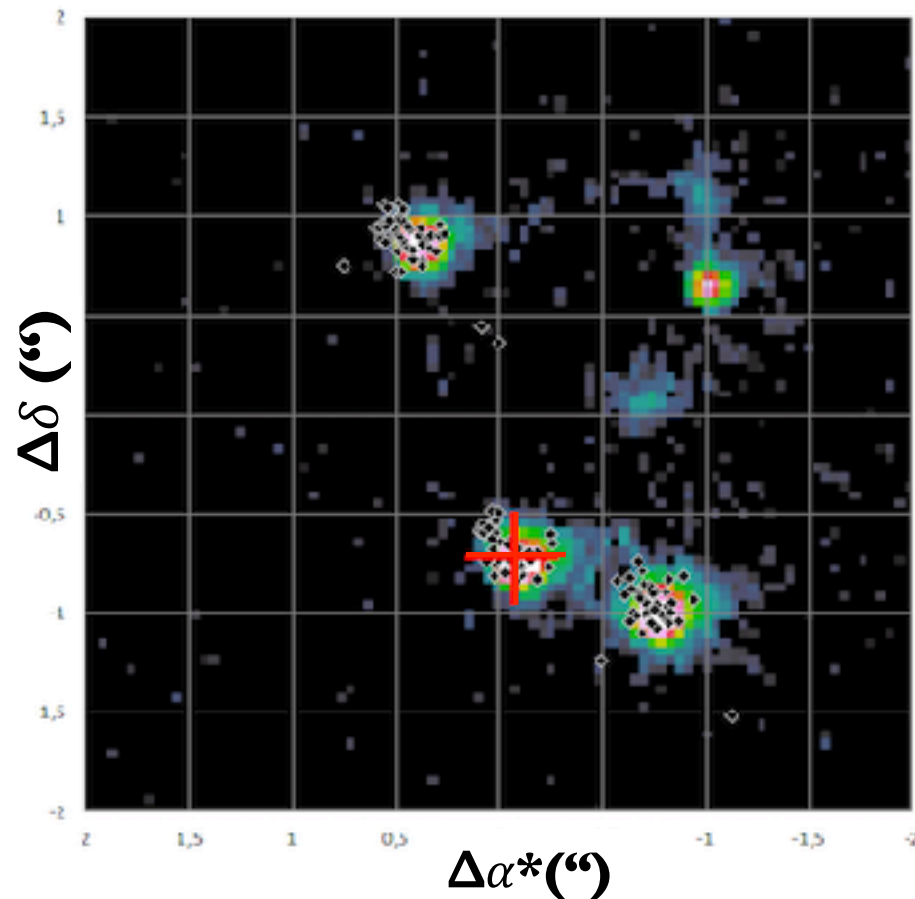
IDT → DR2



B0218+357 (19.28)

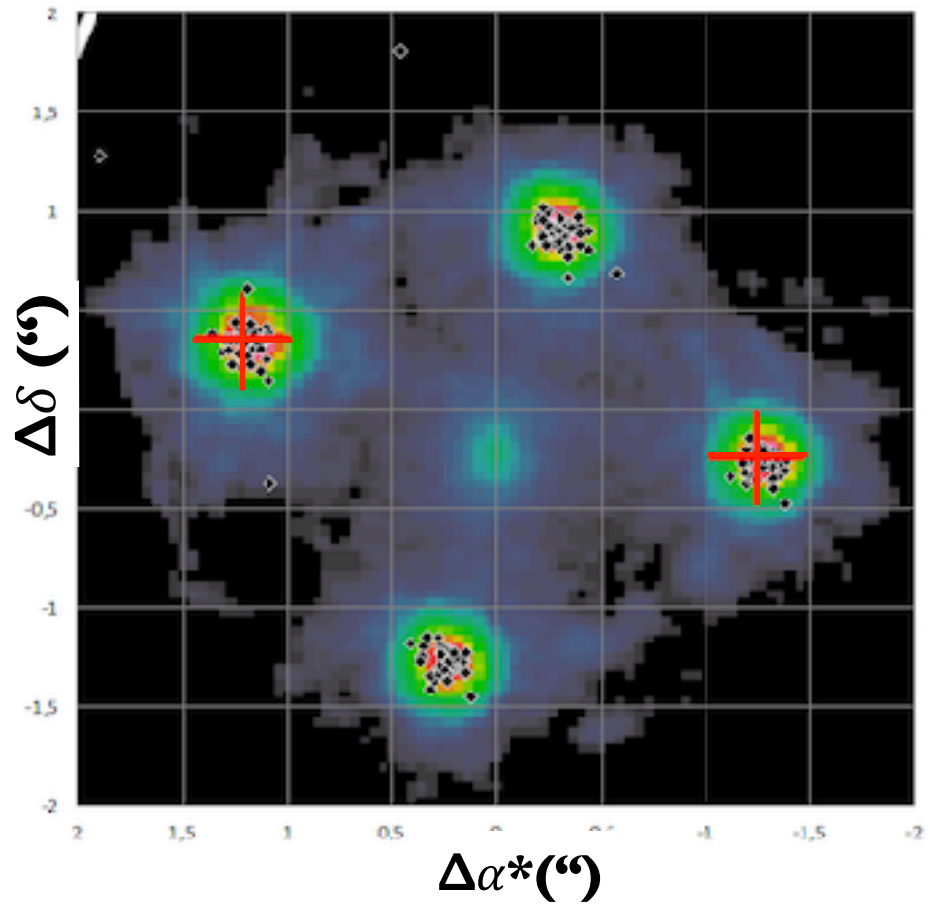


HE0230-2130 (18.00)

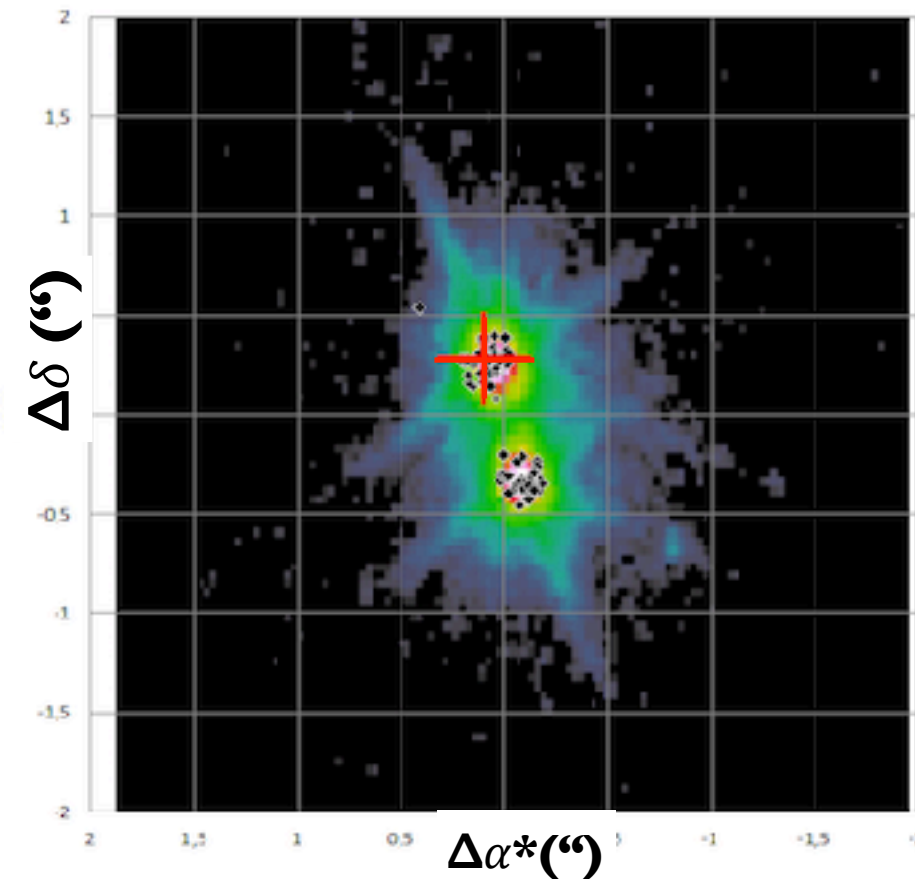


+ DRI
 ■ IDT mes.
 HST images

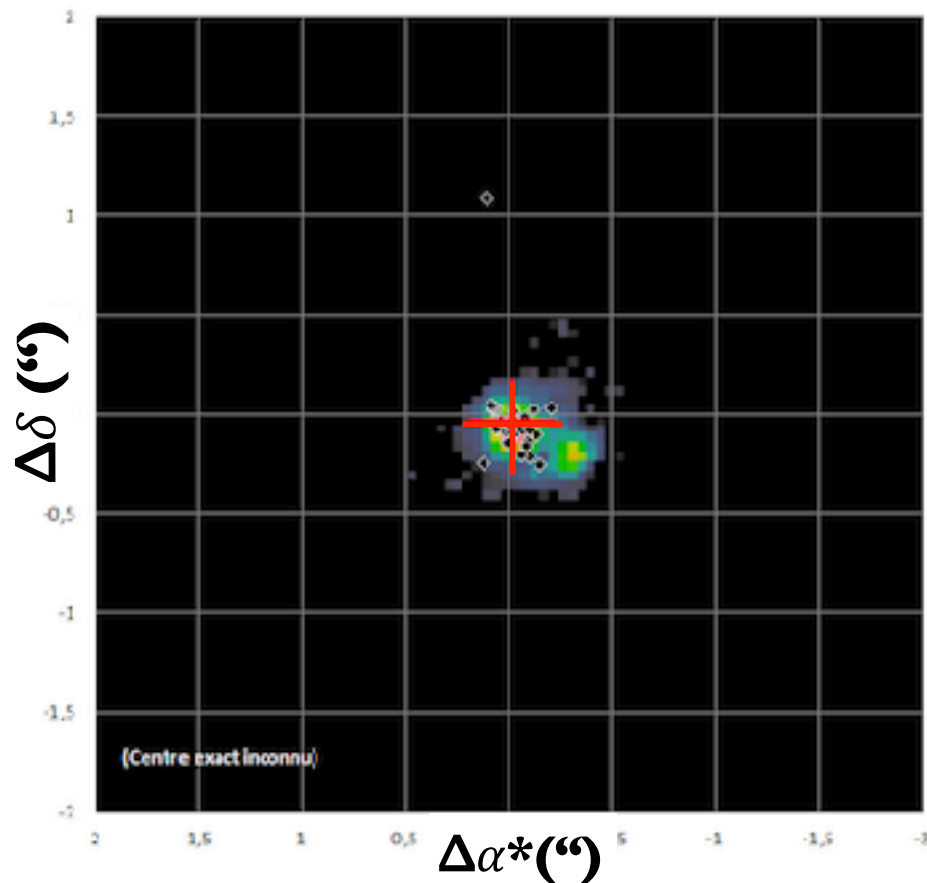
HE0435-1223 (18.05)



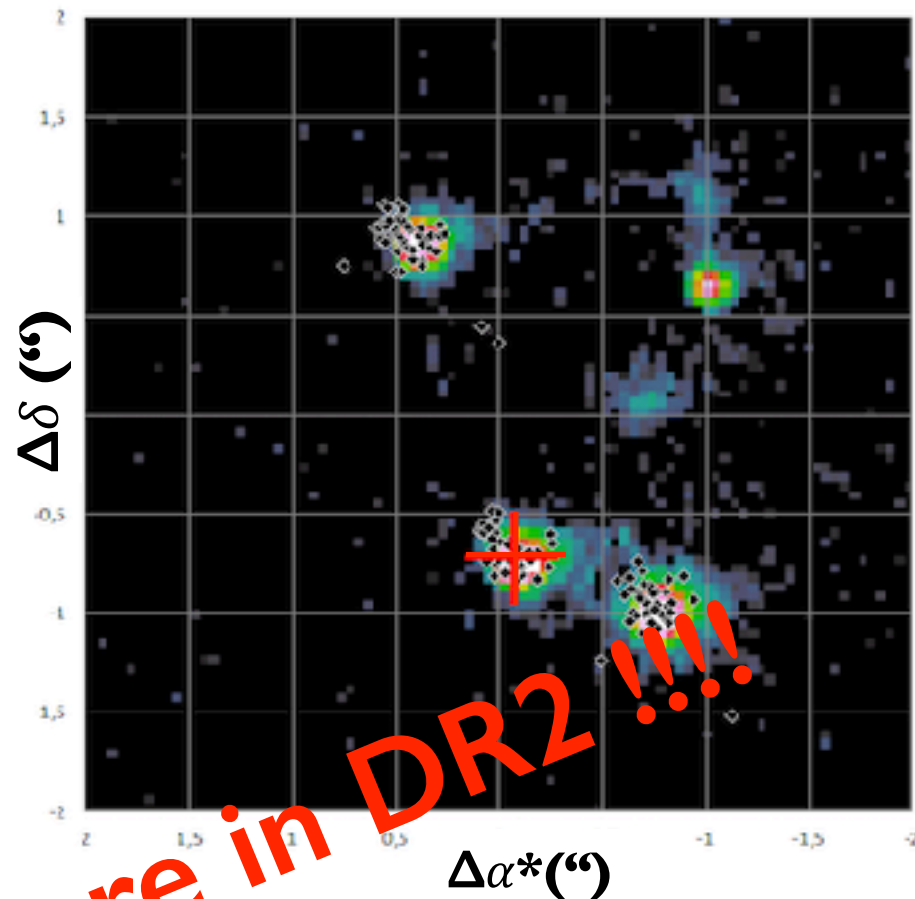
HE0512-3329 (16.27)



B0218+357 (19.28)



HE0230-2130 (18.00)

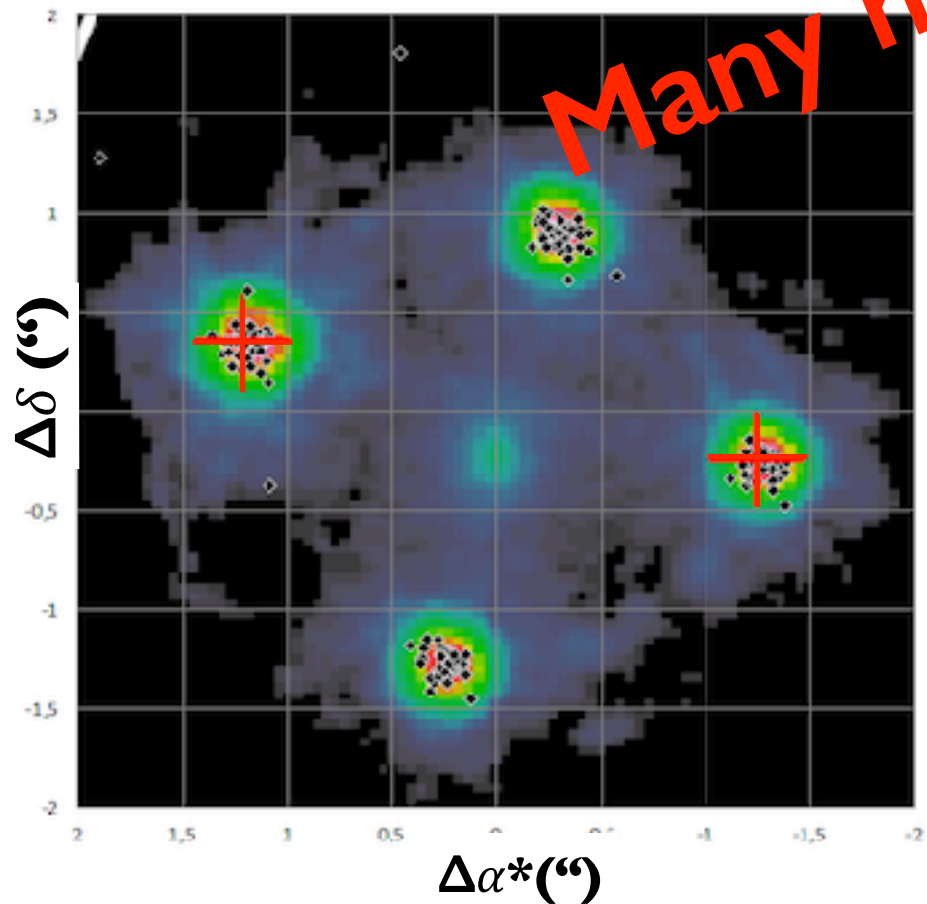


- + DRI
- IDT mes.

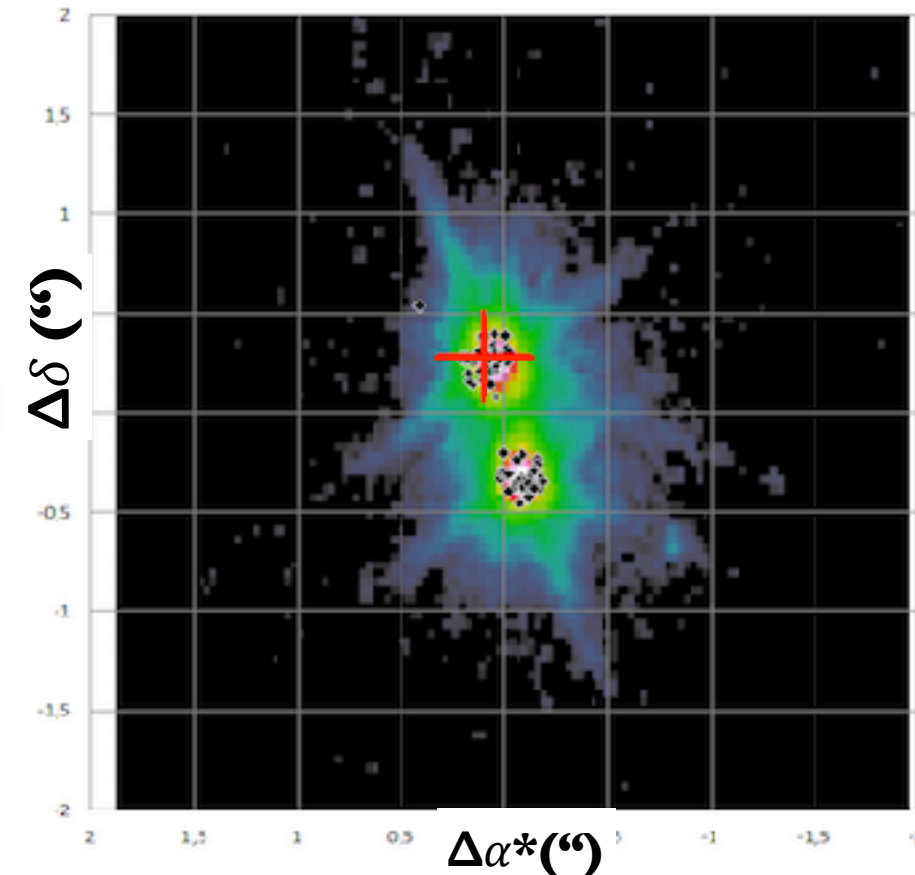
HST images

Many more in DR2 !!!!

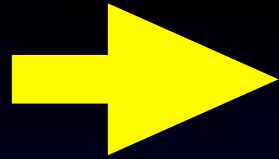
HE0435-1223 (18.05)



HE0512-3329 (16.27)



G magnitude



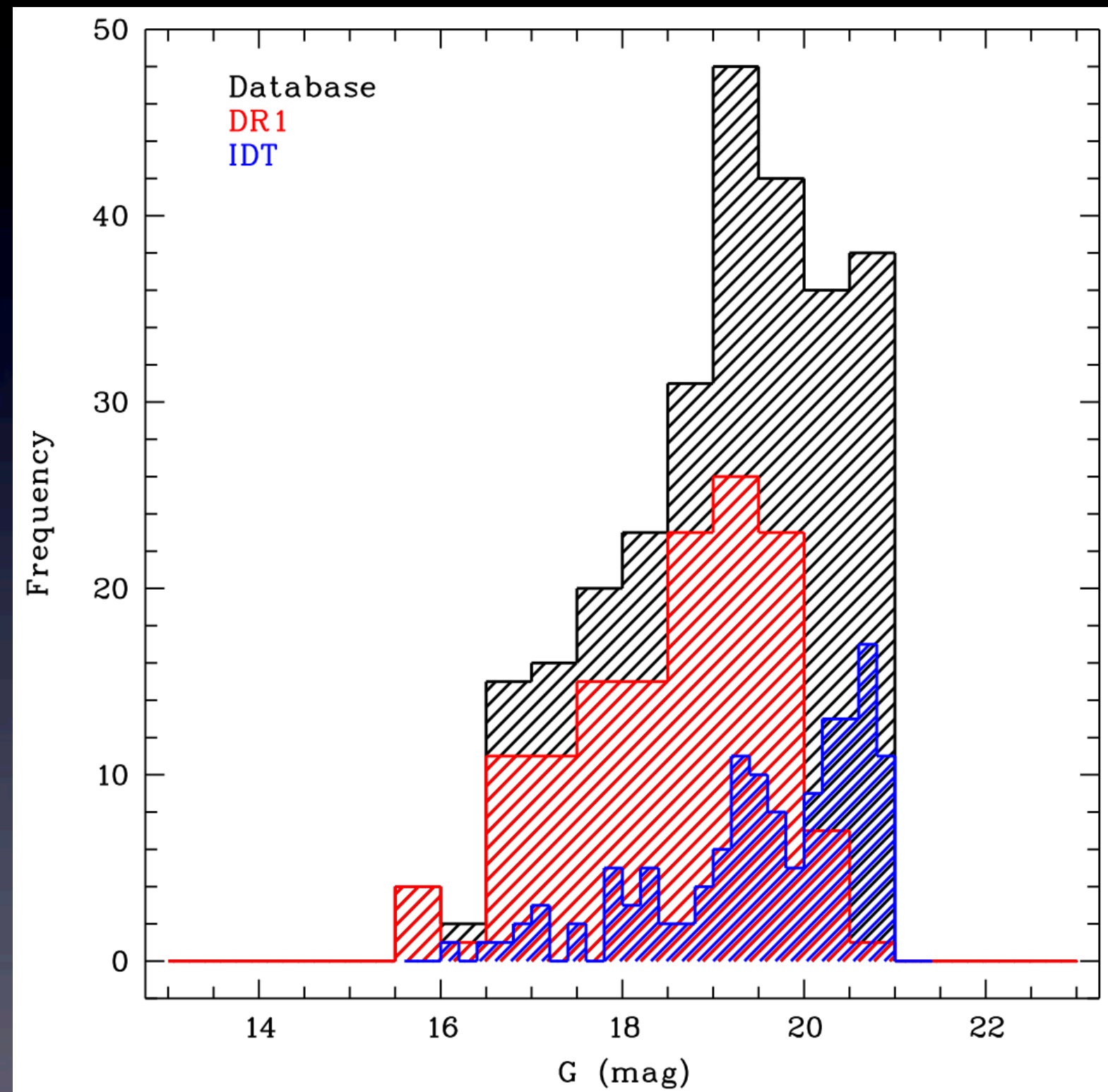
Database : 276 lenses

Detection : 147

(not all components)

- 79 partially in DR1
- 68 only in IDT

- **Reliable astrometry**
- **Expect many more in DR2**

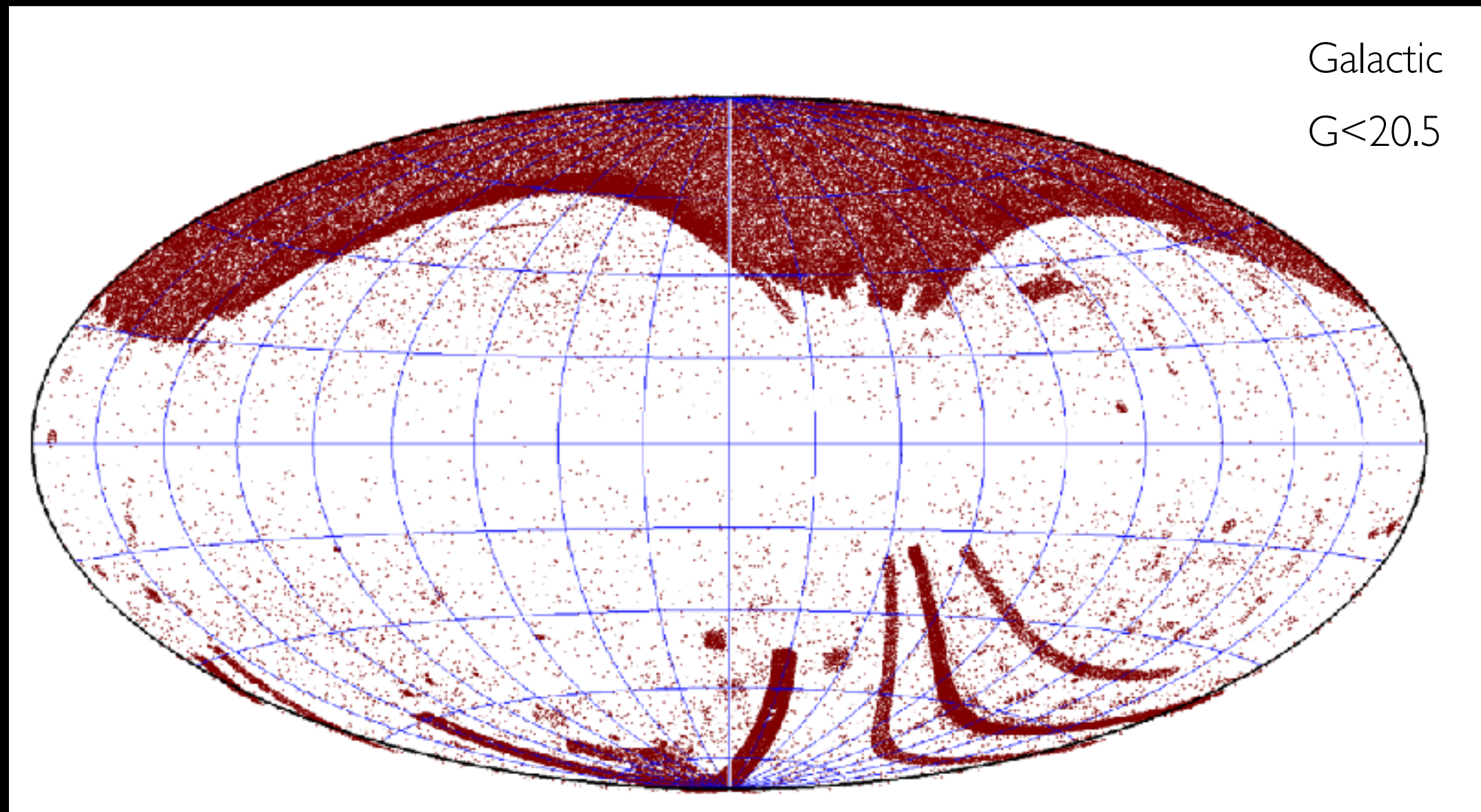


2 - Automatic search around known quasars

➔ **LQAC3 : Large Quasar Astrometric Catalogue**

(Souchay et al. 2015)

- 322 000 entries (Quasars, 14 000 AGNs, 1200 BL Lac)



Credit : F. Mignard, L. Galluccio

Mining and clustering Gaia observations

➔ 1 - Extraction of LQAC3 Quasars observations

- Explore 30 billions of Gaia transits
- 4.2 millions match with a LQAC source ($< 3''$)

➔ **250,000** quasars detected (217 000 with 5+ observations)

➔ 2 - Clustering analysis

- identify groups of points in the area ($\text{rad}=3''$) = sources
- identify cluster of sources (incl. quasar)

➔ unique source, no outlier, clustered 175,000 80%

➔ unique source, $< 20\%$ outliers 32,000 15%

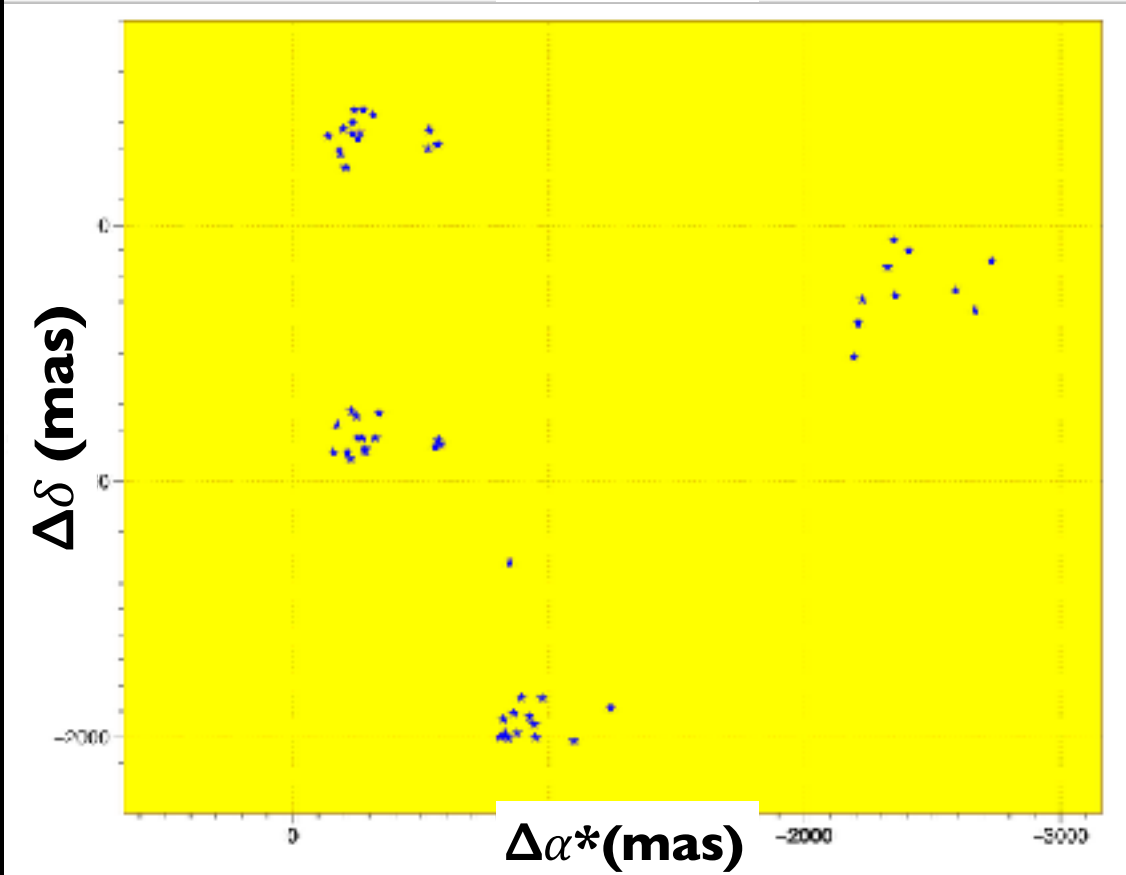
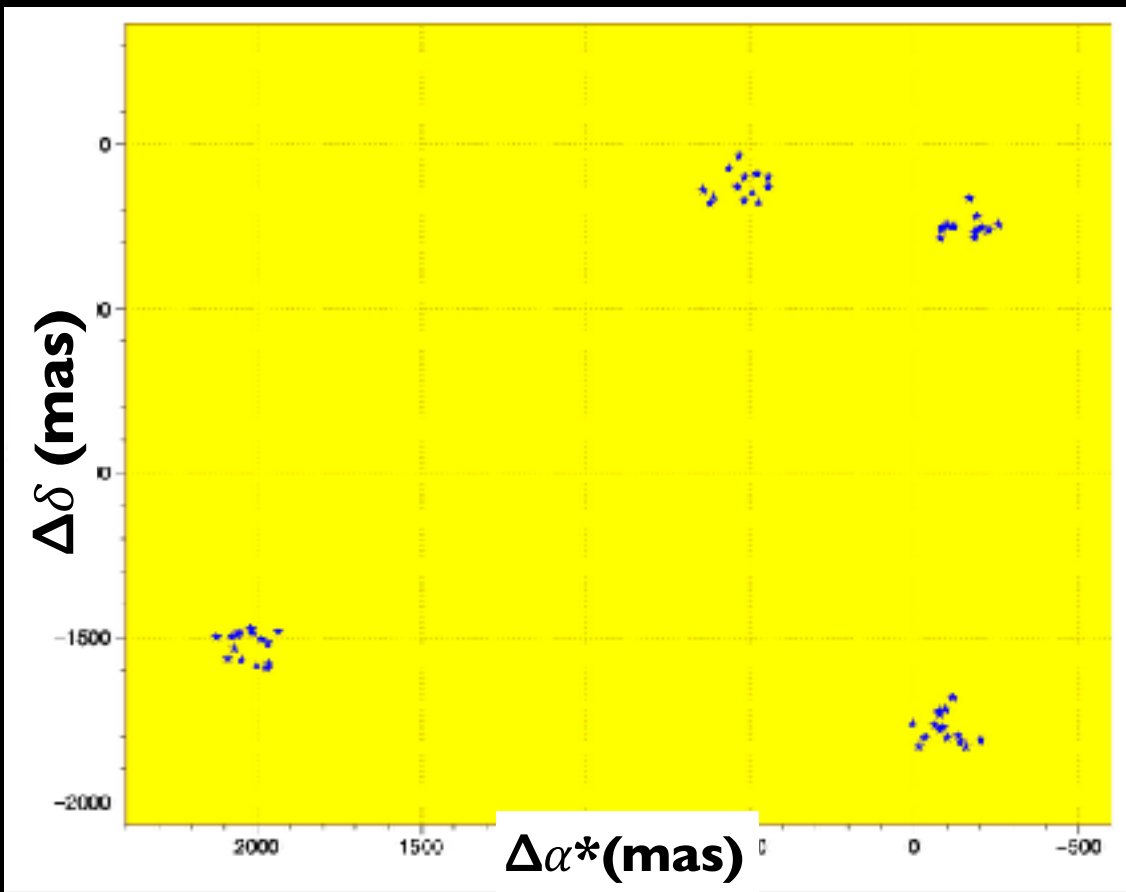
➔ anomalous observations, large scatter ~ 2000

➔ **2 point sources** ~ 1300

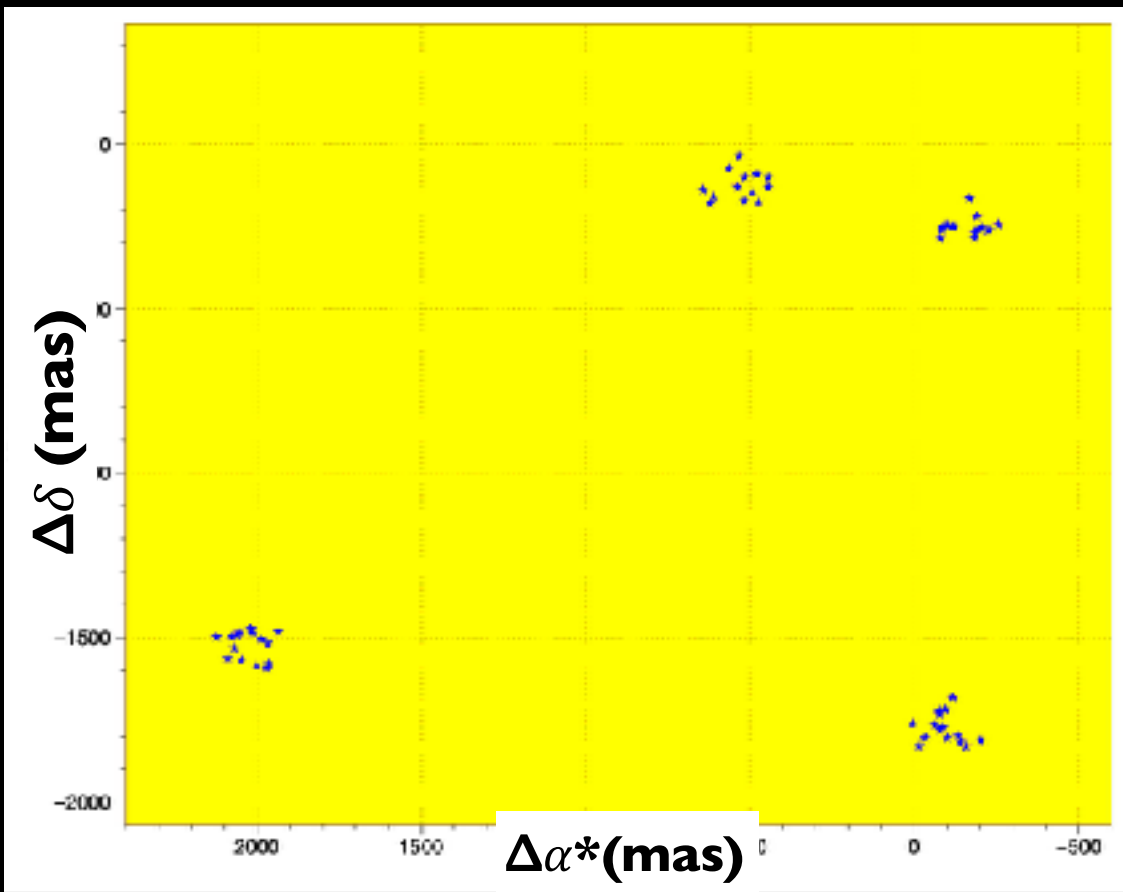
➔ **3 sources** ~ 70

➔ **≥ 4 point sources or images** 10

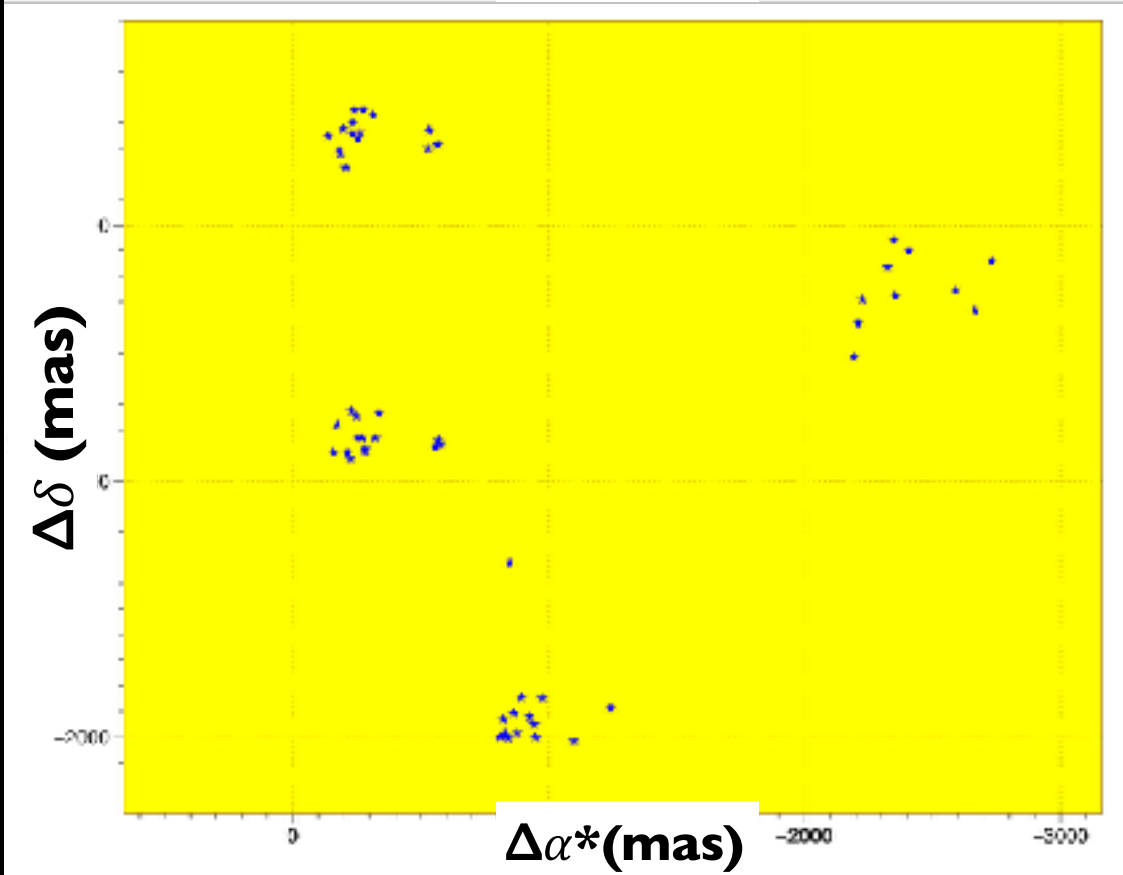
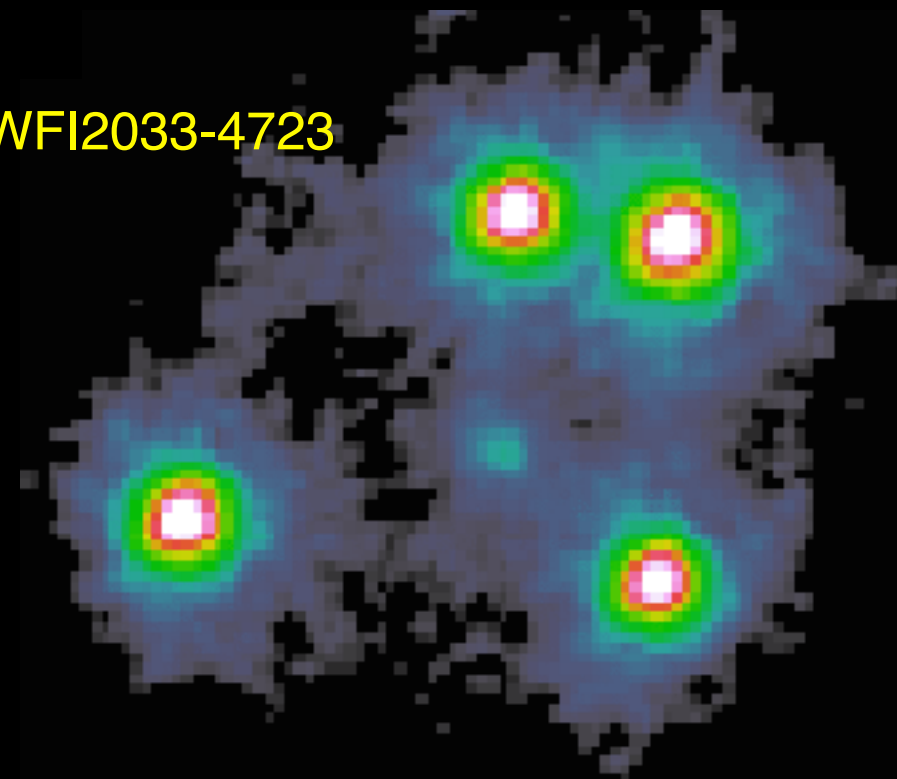
4 images configurations



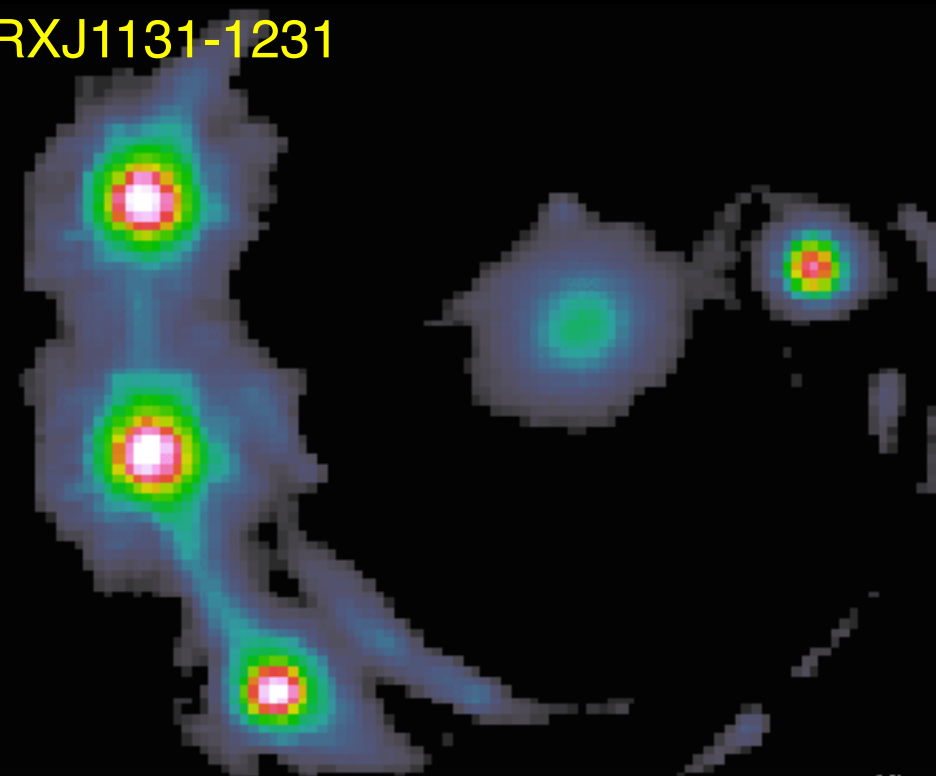
4 images configurations



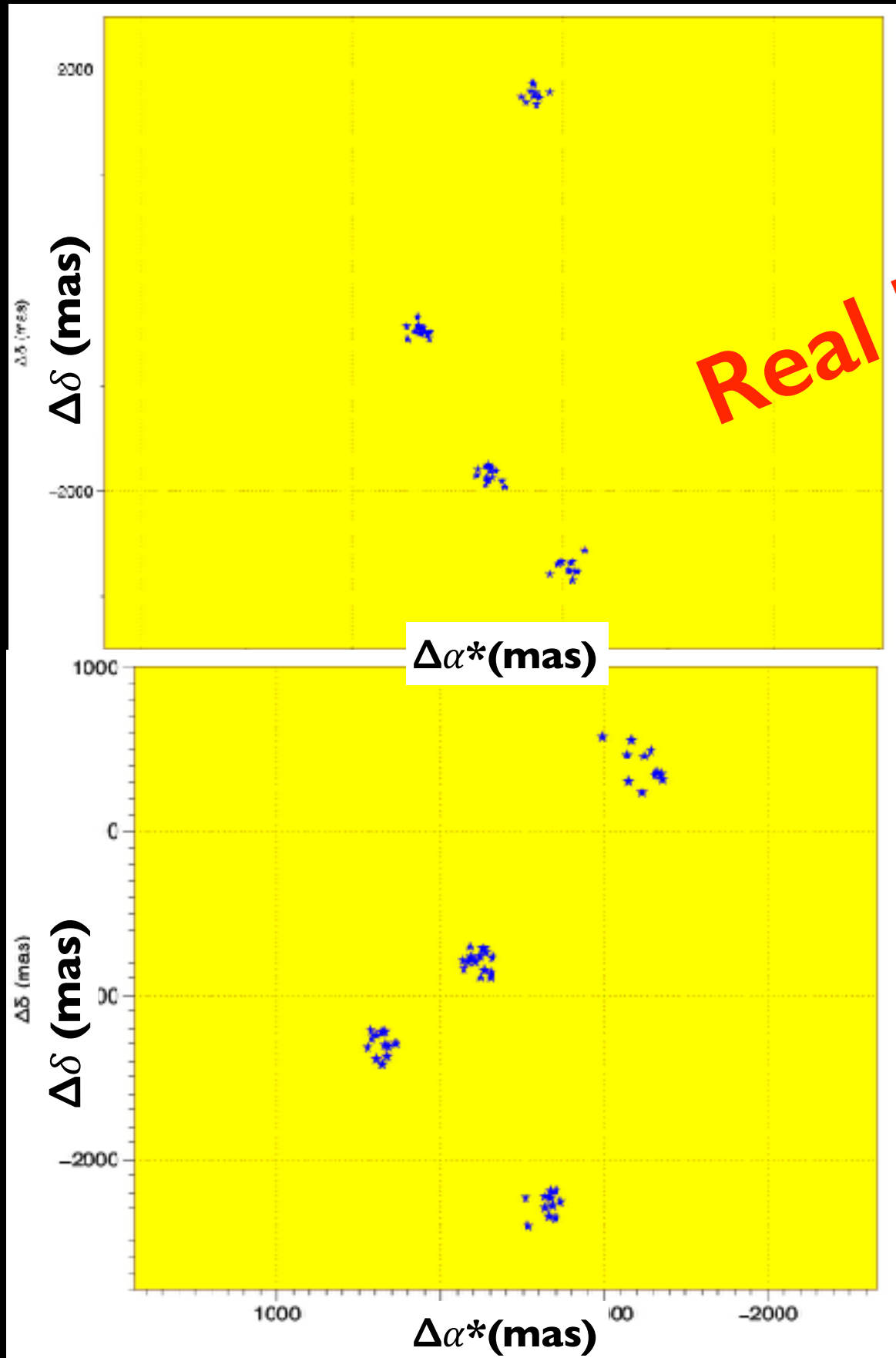
HSWFI2033-4723



RXJ1131-1231



Others 4 images configurations



Real lenses? →

• Good Candidates :

- 29 (3 or 4 components, $<4''$)
- 39 (2 components, $<1''$ high galactic latitude)
- **Necessity to validate the new candidates with :**
 - 1 - Gaia photometry
 - 2 - Lens modelling

➔ **Wait for DR2**

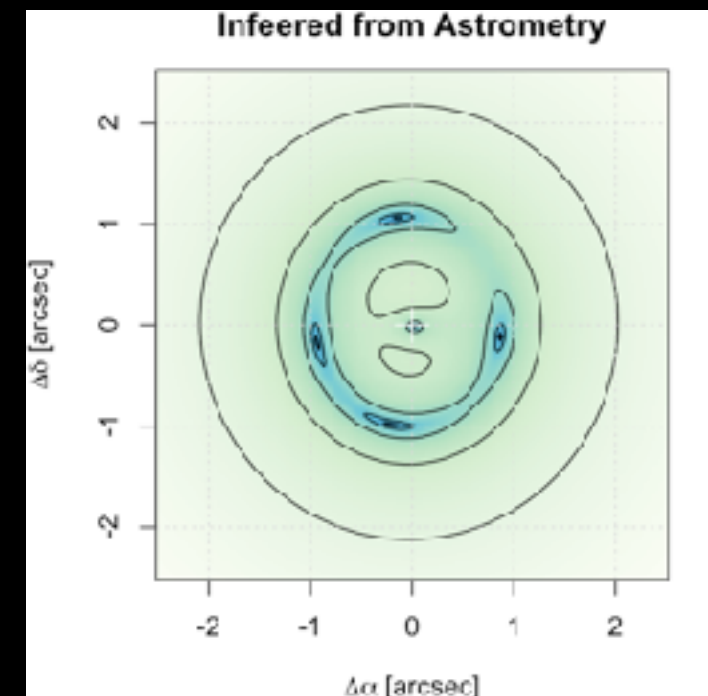
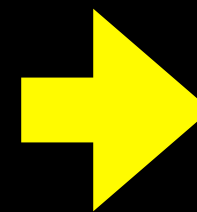
3 - Automatic blind Detection in Gaia DR1

SELenA : Systematic Exploration of Lenses from Astrometry

- An “intelligent” and adaptable framework (kd-trees + machine learning based classifier + lens direct model) for searching lenses in **astrometric** catalogues.
- Can consider *additional* constrains: *photometry, spectroscopy, time-series, ...*



Dedicated UV30 machine in São Paulo
(Gaia DR1 in memory + SSD disks)



- Gaia DR1 = **1.5 G sources**
- Processing time huge = 197 days in a single core ~4 days in 48 cores (*not including lens system modelling!*)

SELenA method

Source
Catalogue

Tree from similarity in
(RA, Dec) space

Leaf similarity check in
other param. spaces
(Gmag, color, pm)

Not a lens

No

yes

SELenA method

Source
Catalogue

Tree from similarity in
(RA, Dec) space

Leaf similarity check in
other param. spaces
(Gmag, color, pm)

Not a lens

No

yes

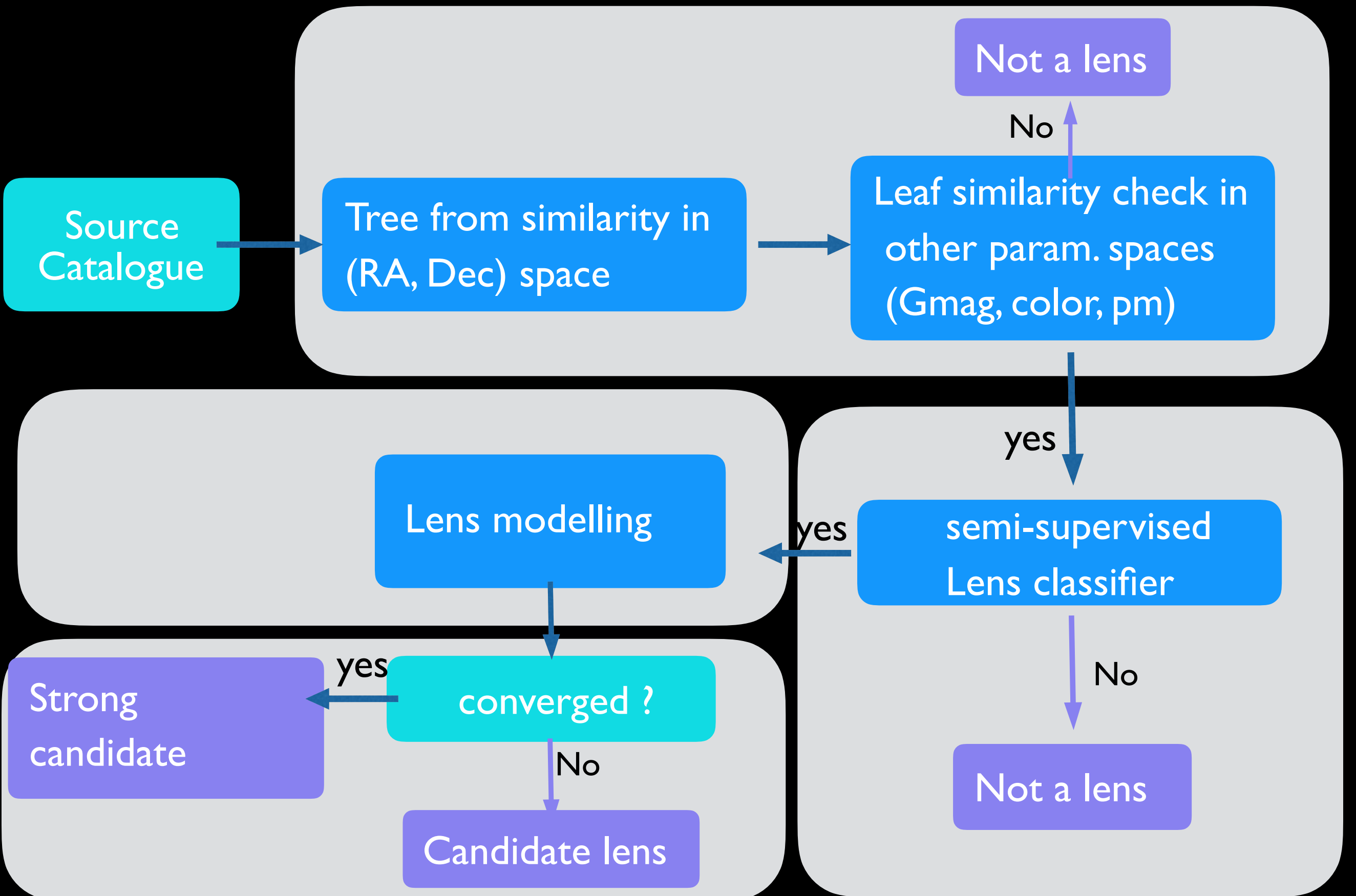
semi supervised
Lens classifier

yes

No

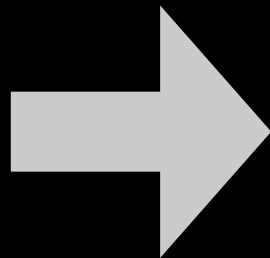
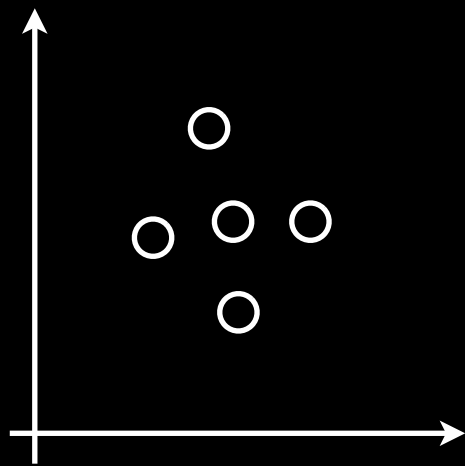
Not a lens

SELenA method



Lens modelling : the problem to be solved

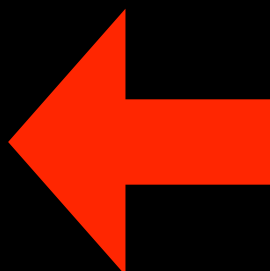
Gaia data



**give-me all the parameters
of the lens system!**

Deflector :: axis ratio, major axis,
Einstein radius ($H_0 + \text{Mass} + \text{redshift}$),
pos. angle, shear terms (g , ang.)
Background QSO :: ang.dist to lens
pos.angle

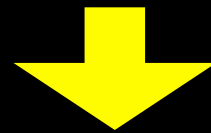
But to do that, we need something that allow us to do the inverse!



Deflector :: axis ratio, major axis,
Einstein radius ($H_0 + \text{Mass} + \text{redshift}$),
pos. angle, shear terms (g , ang.)
Background QSO :: ang.dist to lens
pos.angle

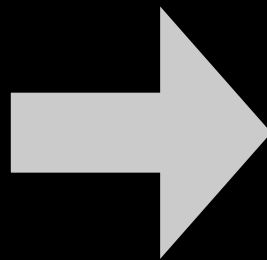
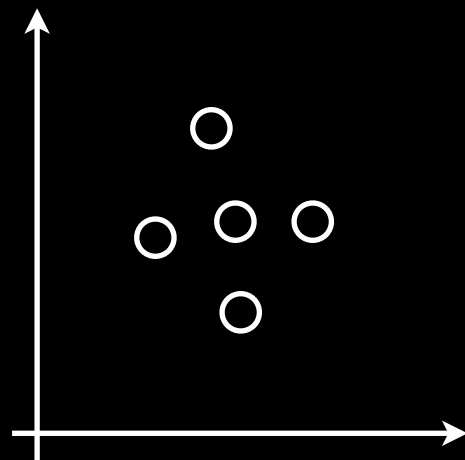
Lens modelling : the problem to be solved

- Forward model implemented in OpenCL (runs in CPU, GPU or FPGA) : **millions of lens configurations can be simulated per second.**

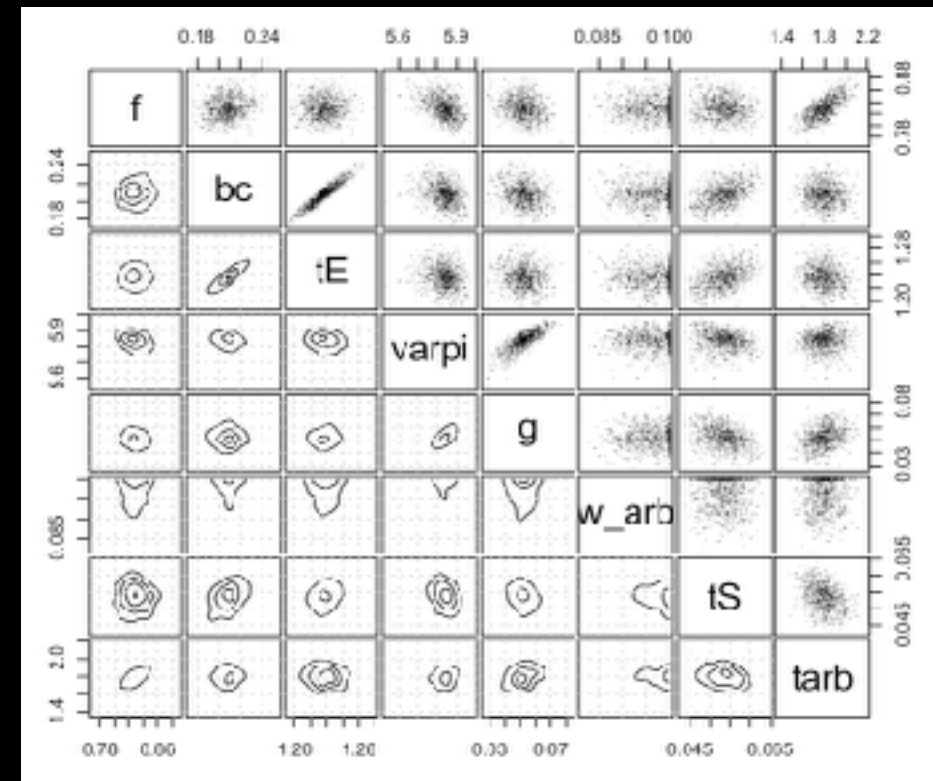


Efficient sampling of the parameter space for the Bayesian inference process

Gaia data



Parameter inference



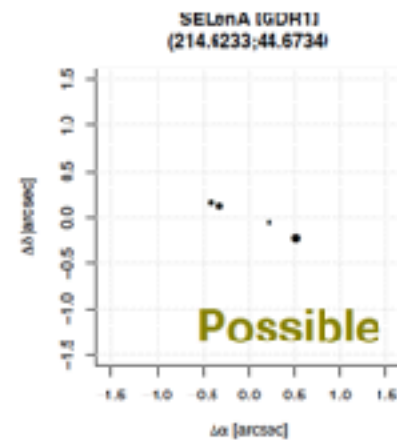
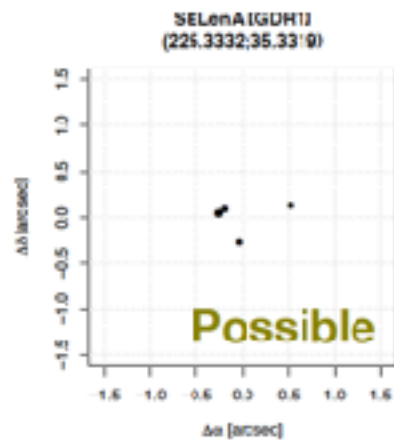
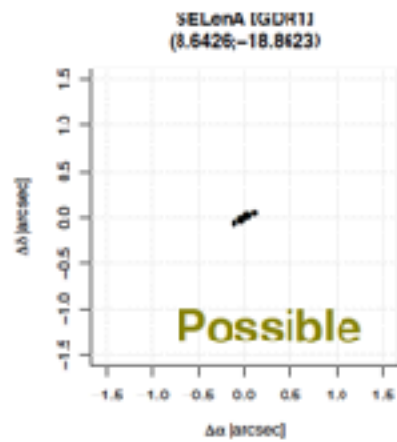
Preliminary Results on DRI

- **Automatic processing outside Galactic Plane** ($D_{\text{pos}} < 1.5''$, $D_{\text{mag}} < 1 \text{ mag}$)

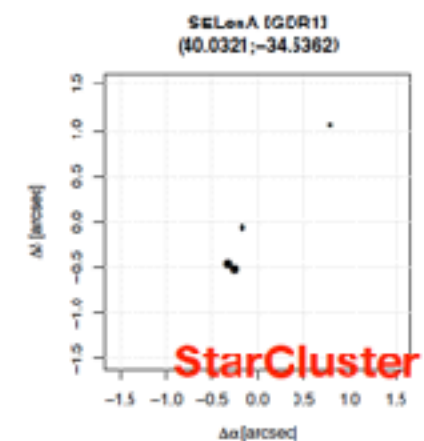
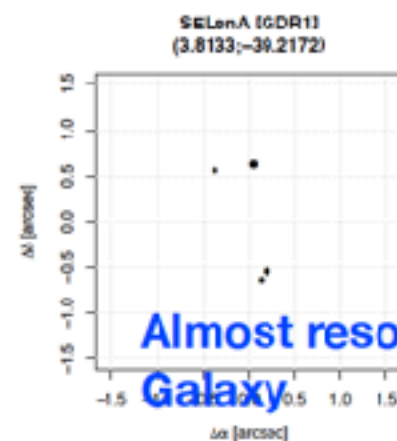
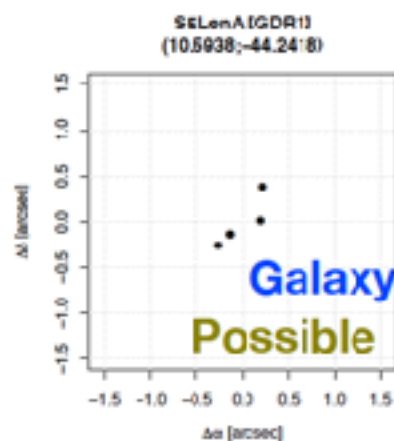
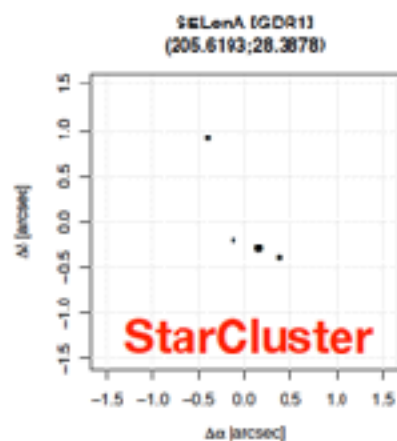
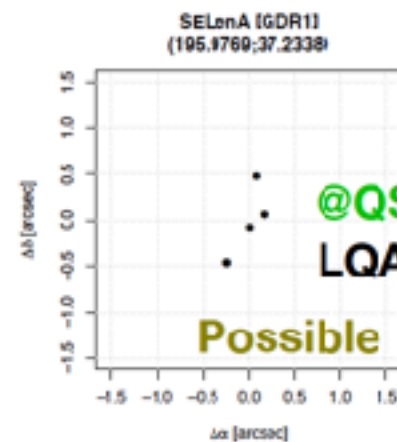
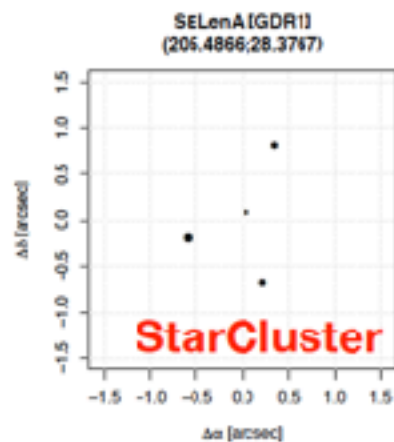
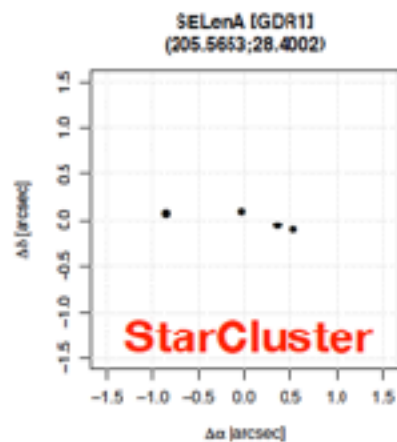
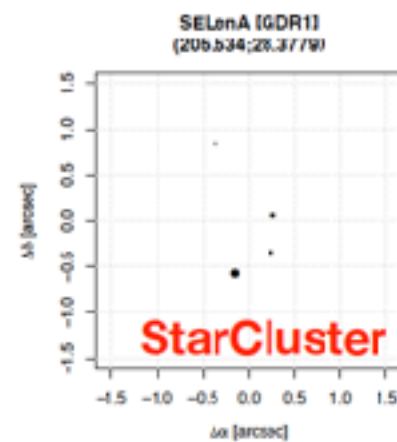
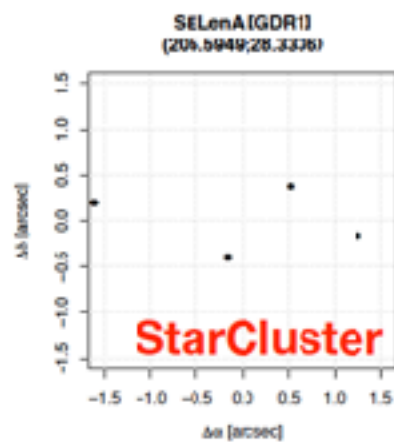
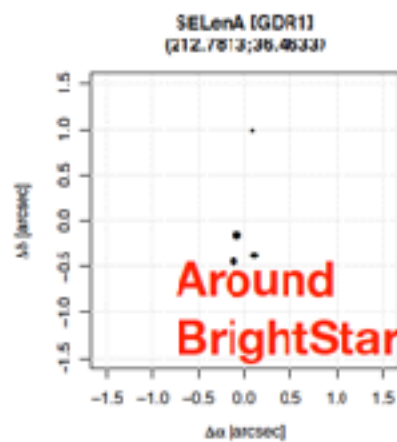
$ b $ cut	Individuals	2-sources	3-sources	4-sources
50° to 55°	6 546 674	10 905	139	0
55° to 60°	5 284 145	9 351	126	0
$> 60^\circ$	13 233 625	35 199	621	13

- **No classifier and lens modelling yet !**

4 images candidates from DR1



Possible
@QSO
Galaxy
StarCluster
BrightStar



Conclusion

- List of good candidates with 3 and 4 images from DR1.
- Proposal at the DOT (Devasthal Optical Telescope) 3.6m telescope (India) for a multi-photometry validation of DR1 candidates.
- We expect many more/complete configurations in DR2.



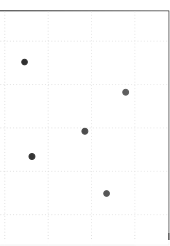
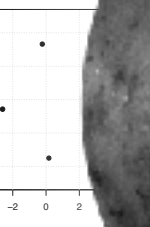
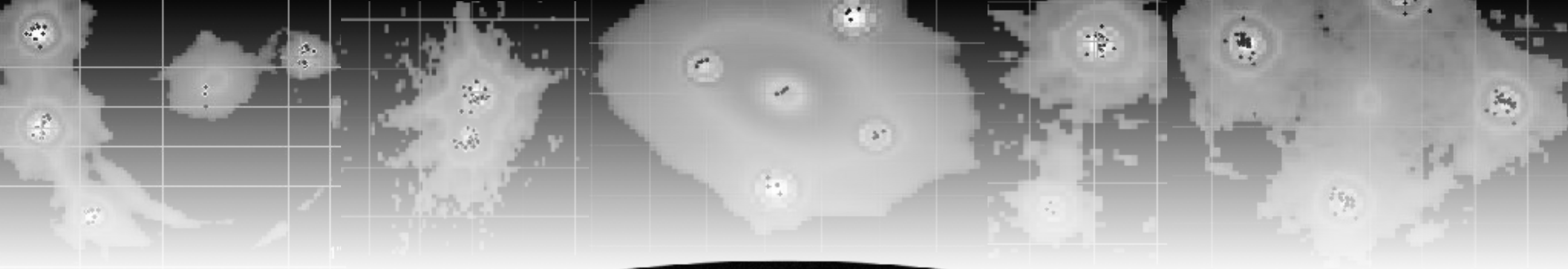
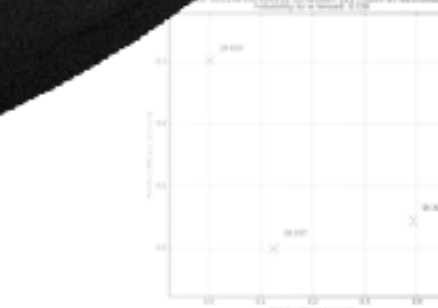
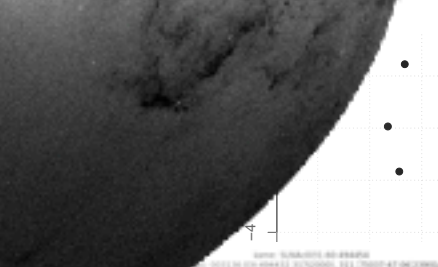
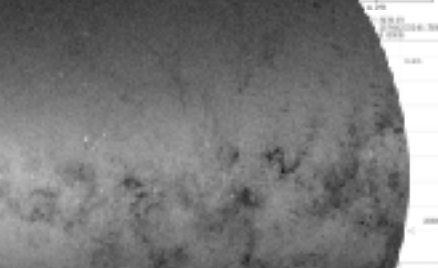
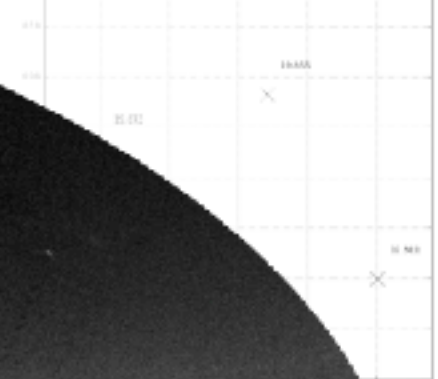


FIGURE 1: SUBSTRUCTURE OF IC 3639
Position: (RA, DEC) = (251.21304, 50.43345) (J2000.0)
Threshold to be fitted: 8.950



Thank you!

